VLA OBSERVING APPLICATION

A	
rcvd:	

DEADLINES: 1st of Feb., June., Oct. for next configuration following review INSTRUCTIONS: Each numbered item must have an entry or N/A E-MAIL TO: propsoc@nrao.edu (different for some Rapid Response Science)

OR MAIL TO: Director NRAO, 520 Edgemont Rd., Charlottesville, VA 22903-2475

(1) Date Prepared:

(2) Title of Proposal: A second maser in IC 10

				Students Only		
(3) AUTHORS	INSTITUTION	E-mail	G/U	For	Ph.D.	
(Add * for new location)				Thesis?	Year	
A. Brunthaler	JIVE, Dwingeloo	brunthaler@jive.nl				
H. Falcke	ASTRON, Dwingeloo	falcke@astron.nl				
M.J. Reid	CfA Harvard/Smithsonian	reid@cfa.harvard.edu				
L.J. Greenhill	CfA Harvard/Smithsonian	lgreenhill@cfa.harvard.edu				
C. Henkel	MPIfR, Bonn	p220hen@mpifr-bonn.mpg.de				

C. Helikel	MII III, DOIIII		pzzonen@mpm-i	bonn.mpg.de			
(4) Related VLA previous p	proposal number(s)	:					
(5) Contact author			(6) Telephone:	+31 521 596	540		
` '	Brunthaler		E-mail:				
9	nt Institute for VLI	BI in Europe	Fax:	v			
	tbus 2	•					
799	0 AA Dwingeloo, N	Vetherlands					
(7) Scientific Category: (solar system (galactic 🛇	extragalactic (O other:			
()	, ,	•	9	<u> </u>			
Rapid Response Science	_	ansient (X) I	Exploratory	O Target of Op	portunity		
(8) Configurations (one per	,						
(A+Pt, A, B, C, D, BnA, C	CnB, DnC, Any)	В					
(9) Wavelength(s)		1.3 cm					
(400, 90, 20, 6, 3.5, 2, 1)	1.3, 0.7 cm						
(10) Time requested							
(hours)		0.5 hours					
(11) Type of observation: (check all that apply)	_	⊗ spectroscop high-time resol		nnel continuum own link	Opolarimet	ry Os	solar
(12) Suitable for dynamic s	scheduling? \otimes S	Suitable	O Unsuitable				
(13) ABSTRACT (do not v	write outside this s	pace)					
We propose to obtain an accin our ongoing project to n density of our current target the other hand, IC 10-NW h has also a broader spectrum. We would like to get the pobe correlated.	neasure the proper source in IC 10 (IC as been more stable than IC 10-SE, ma	motions of the 210-SE) has dec e since its disco aking it already	e two Local Grou creased over the l every in the late 1 v a better target	up galaxies IC 10 ast years from > .980s at a level o source for our V	0 and M33. The 1 Jy to ~ 0.1 Jy from IC 1 LBA observation	ne flux Jy. On 0-NW ons.	

NRAO use only (08/03)

(14) Observer present for observations:	es W No Data	allalysis at: W 1	ionie O AOC oi	C V (2 weeks notion
(15) Help required: \bigcirc None \bigcirc Consul	tation Friend	(extensive help)		
(16) Spectroscopy only	line 1	line 2	line 3	line 4
Transition (HI, OH, etc.)	H ₂ O			
Rest Frequency (MHz)	22.235			
Velocity (km/s)	-330			
Observing frequency (MHz)				
Correlator mode	1A			
IF bandwidth(s) (MHz)	6.25			
Hanning smoothing (y/n)	n			
Number of channels per IF	128			
Frequency Resolution (kHz/channel)	48.828			
Rms noise (mJy/bm, nat. weight., 1 hr)	2.5			
Rms noise (K. nat. weight., 1 hr)				

(17) Number of sources:	1	

(If more than 10 please attach list. If more than 30 give only selection criteria and LST range(s).)

range	(hr)
>20	0.5

^{*}For spectral line, this should be the total flux at the peak of the line

Notes to the table (if any):

- (19) Restrictions to elevation (other than hardware limits) or HA range (give reason):
- (20) Preferred range of dates for scheduling (give reason):
- (21) Dates which are not acceptable:
- (22) Special hardware, software, or operating requirements:
- (23) Please attach a self-contained <u>Scientific Justification</u> not in excess of 1000 words. (Preprints or reprints will be ignored.) Please include the full addresses (postal and e-mail) for first-time users or for those that have moved (if not contact author). When your proposal is scheduled, the contents of the cover sheets become public information (Any supporting pages are for refereeing only).

v4.2 8/03