VLA OBSERVING APPLICATION



DEADLINES: 1st of Feb., June., Oct. for next configuration following review INSTRUCTIONS: Each numbered item must have an entry or N/A E-MAIL TO: propsoc@nrao.edu (different for some Rapid Response Science)

OR MAIL TO: Director NRAO, 520 Edgemont Rd., Charlottesville, VA 22903-2475

(1) Date Prepared: Aug. 10, 2004

(2) Title of Proposal: Exploratory H₂O Maser Observations Toward W33A

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				Students Only		
(3) AUTHORS	INSTITUTION	E-	E-mail			Ph.D.
(Add * for new location	on)			Thesis?	Year	
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(4) Rolated VI A prov	vious proposal number(s):					
(4) Related VLA pre-	vious proposai numbei (s).					
(5) Contact author		(6) Telephone:	808-932-2319			
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(4) Related VLA previou	$_{ m ls}$ proposal number(s)	:					
address: I	Crystal Brogan nstitute for Astronom 40 North A'ohoku Pl Iilo, HI 96781	· ·	(6) Telephone: E-mail: Fax:		i.edu		
(7) Scientific Category: Rapid Response Scie	-	•		○ other: ○ Target of Opportu	$_{ m inity}$		
(8) Configurations (one p (A+Pt, A, B, C, D, BnA		A					
(9) Wavelength(s) (400, 90, 20, 6, 3.5, 2	2, 1.3, 0.7 cm)	1.3cm					
(10) Time requested (hours)		2 hr					
(11) Type of observation	continuum	O -		nnel continuum	•	etry C) solar

(TT)	Type of observation.	Continuum	. Specific	scopy C	mundichanner ce	muni	O polarii.	IE
	(check all that apply)	\bigcirc pulsar	O high-time	resolution	O Pie Town lin	nk 🔘 o	ther:	
(12)	Suitable for dynamic s	cheduling?	\bigotimes Suitable	\bigcirc Uns	suitable			

(13) ABSTRACT (do not write outside this space)

We have recently (July 21, 04) obtained high resolution 342 GHz SMA spectral line and continuum data for the high mass protostellar object W33A. From these data combined with existing near infrared data we have been able to trace the location of a possible disk and small scale jet/outflow. The relationship between CH₃OH, OH, and H₂O masers within such regions is currently a matter of vigorous study. In particular we would like to determine whether the H₂O masers trace the disk, jet, or outflow components toward W33A. We have analyzed archival C configuration 1.3 cm data and find evidence that the masers are associated with all three components, but these data are of poor quality and have insufficient spatial resolution to constrain the morphology of the individual masar spots. Additionally the 40 km s¹ bandwidth of these data may not trace all of the maser features. We request 2 hours of exploratory A-configuration time to map out the H₂O masers toward W33A with high resolution and wider velocity coverage.

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(If more than 10 please attach list. If more than 30 give only selection criteria and LST range(s).)

(18) NAME	Coordi: 1950 () RA hh mm	$2000 \bigotimes$ Dec. $\pm xx.x^{\circ}$	Conf.	λ (cm)	Corr. mode	(MHz)	Total Flux (Jy)*	LAS	Required rms (mJy/bm)	Required dynamic range	Time request (hr)
W33A	18 14, -17	7.9	A	1.3	2IF	3.125	3 Ју	0.1"	10	400	2

^{*}For spectral line, this should be the total flux at the peak of the line

Notes to the table (if any):

- (19) Restrictions to elevation (other than hardware limits) or HA range (give reason):
- (20) Preferred range of dates for scheduling (give reason):

Toward the end of the move from D to A configuration would be perfect.

(21) Dates which are not acceptable:

Not during beginning of move from D to A. We need good imaging capabilities and all the resolution we can get.

- (22) Special hardware, software, or operating requirements:
- (23) Please attach a self-contained Scientific Justification not in excess of 1000 words. (Preprints or reprints will be ignored.)

Please include the full addresses (postal and e-mail) for first-time users or for those that have moved (if not contact author).

When your proposal is scheduled, the contents of the cover sheets become public information (Any supporting pages are for refereeing only).

 $v4.2 \ 8/03$